

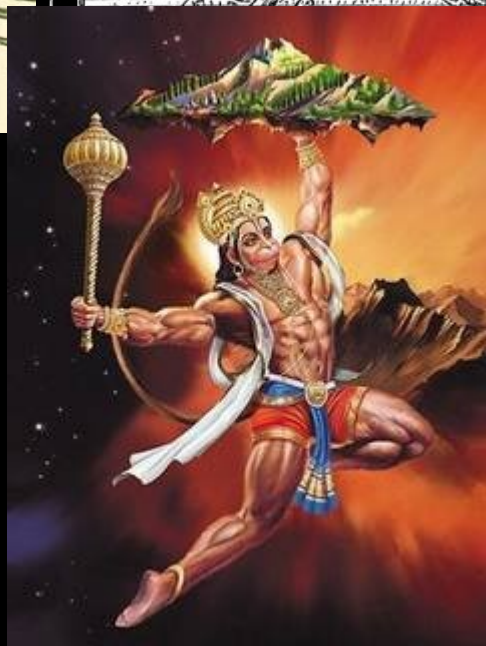
About Mars, its atmosphere and the dust



overview

- * Mars in ancient culture and now
- * Exploration of the Planet
 - missions
 - some discoveries
- * Martian atmosphere
- * Dust in Martian atmosphere, what we really know about it?
- * Scattering on Martian dust
- * Scattering matrix (of Martian dust analogs)
- * Some results
- * Some conclusions

Mars in ancient culture and now



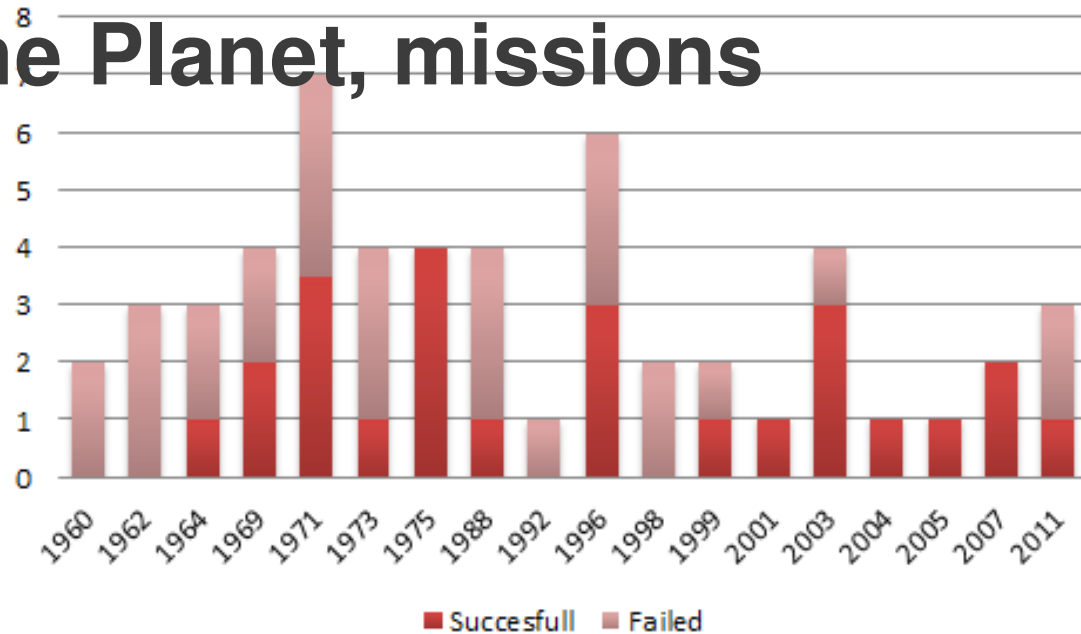
Mars in ancient culture and now



Exploration of the Planet, missions

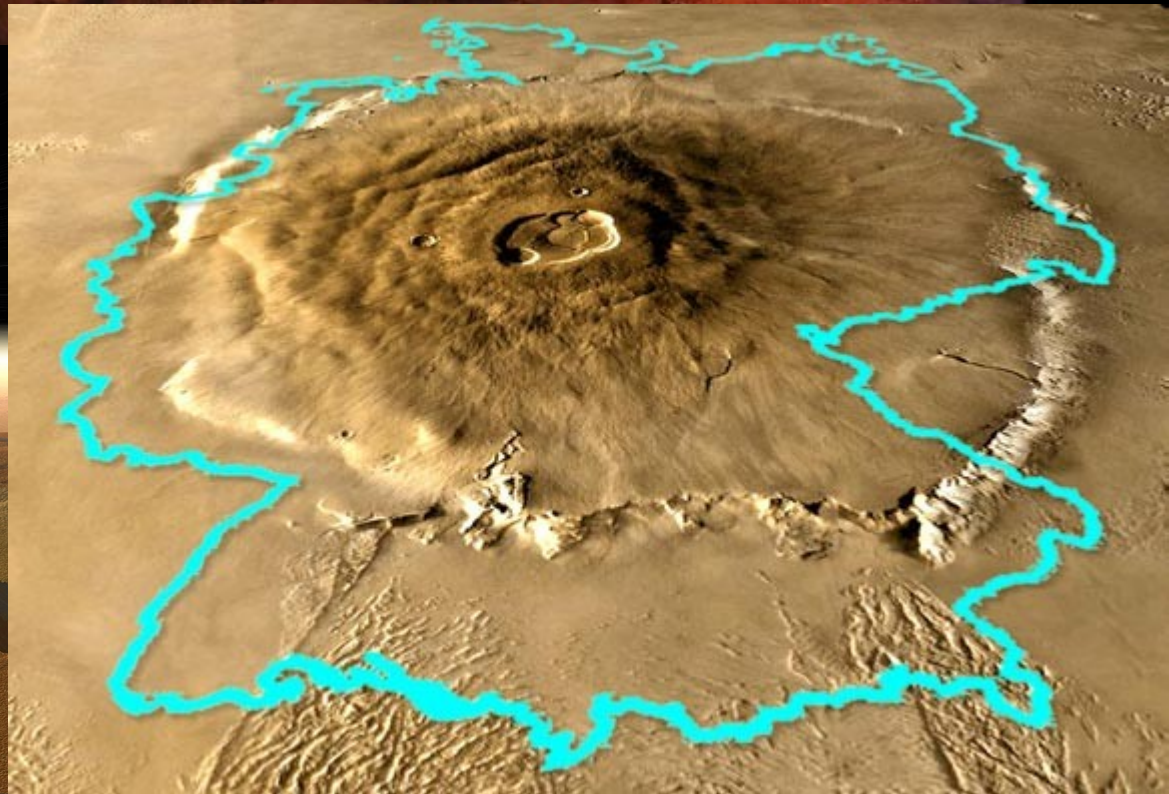
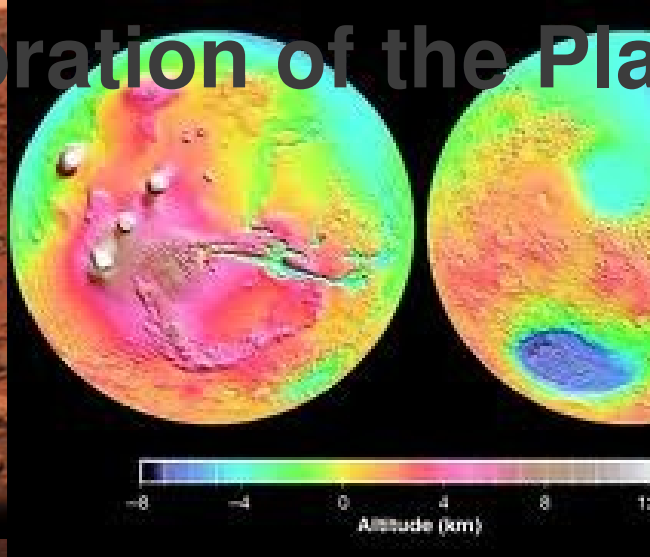
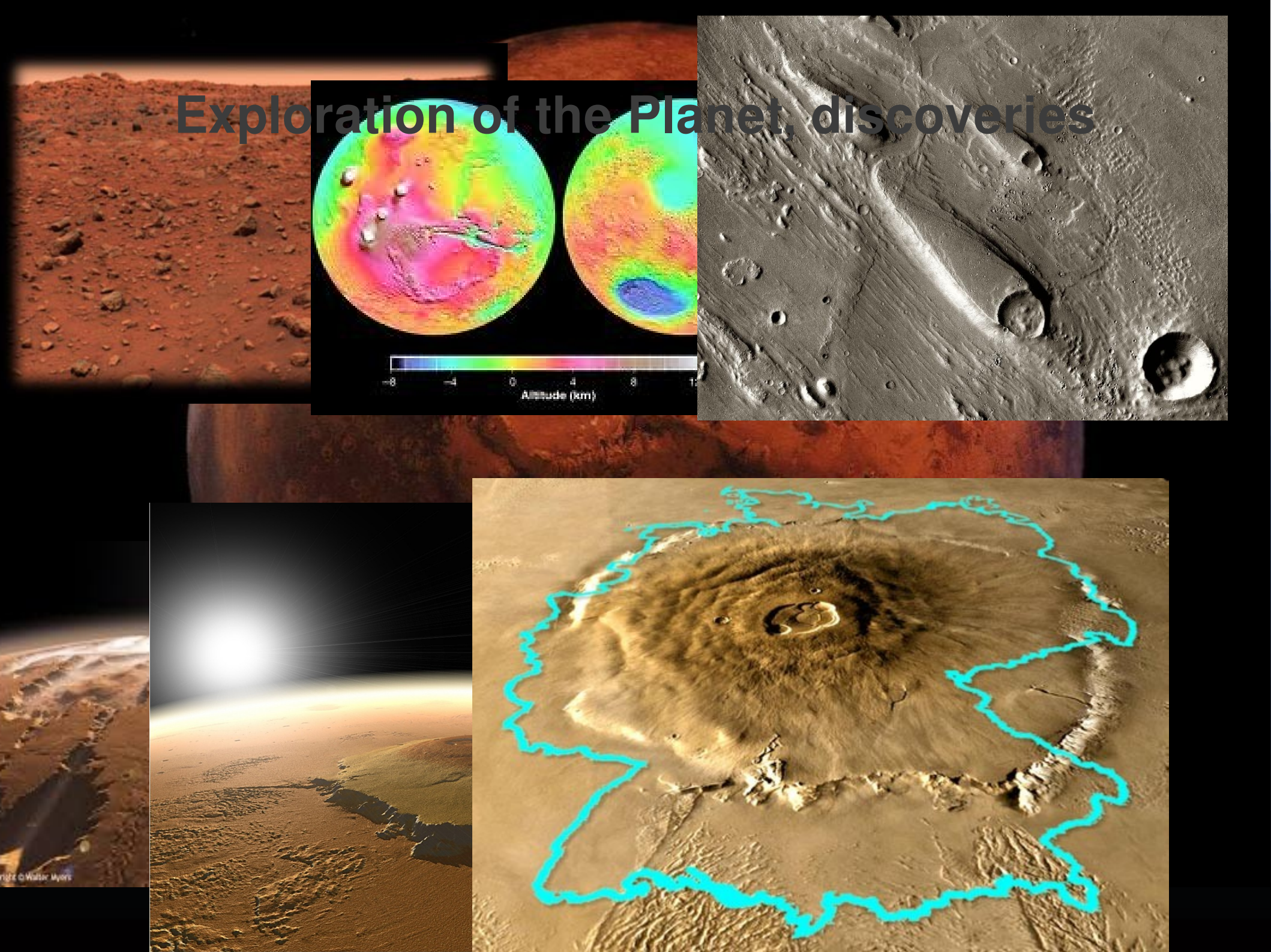


Extensive knowledge



Launch/ /status	Mission	Nation/ /agency
1963	Mars 1	USSR
1964	Mariner 4	NASA
1969	Mariner 6	NASA
1969	Mariner 7	NASA
1974	Mariner 9	NASA
1975	NASA	Viking 1 Orbiter/Lander
1975	NASA	Viking 2 Orbiter/Lander
1988	Phobos-2	USSR
1996 mission ended 2006	Mars Global Surveyor (MGS)	NASA
1996 mission ended 1997	Mars Pathfinder (MPF)	NASA
2001-current	Mars Odyssey Orbiter	NASA
2003-current	Mars Express Orbiter	ESA
2003-current (Opportunity)	Mars Exploration Rovers (MER)	NASA
2003-2011 (Spirit)		
2005-current	Phoenix	NASA
2011-current	2008!!!!	NASA
2013-On way to Mars	Mangalyaan	India
2013-On way to Mars	MAVER	NASA

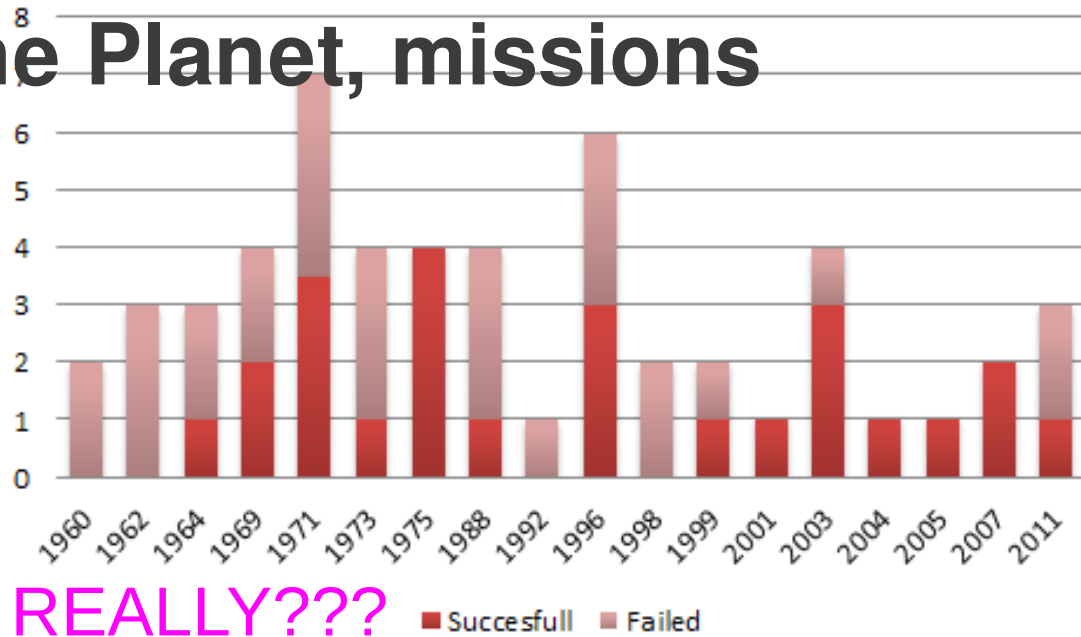
Exploration of the Planet, discoveries



Exploration of the Planet, missions



Extensive knowledge

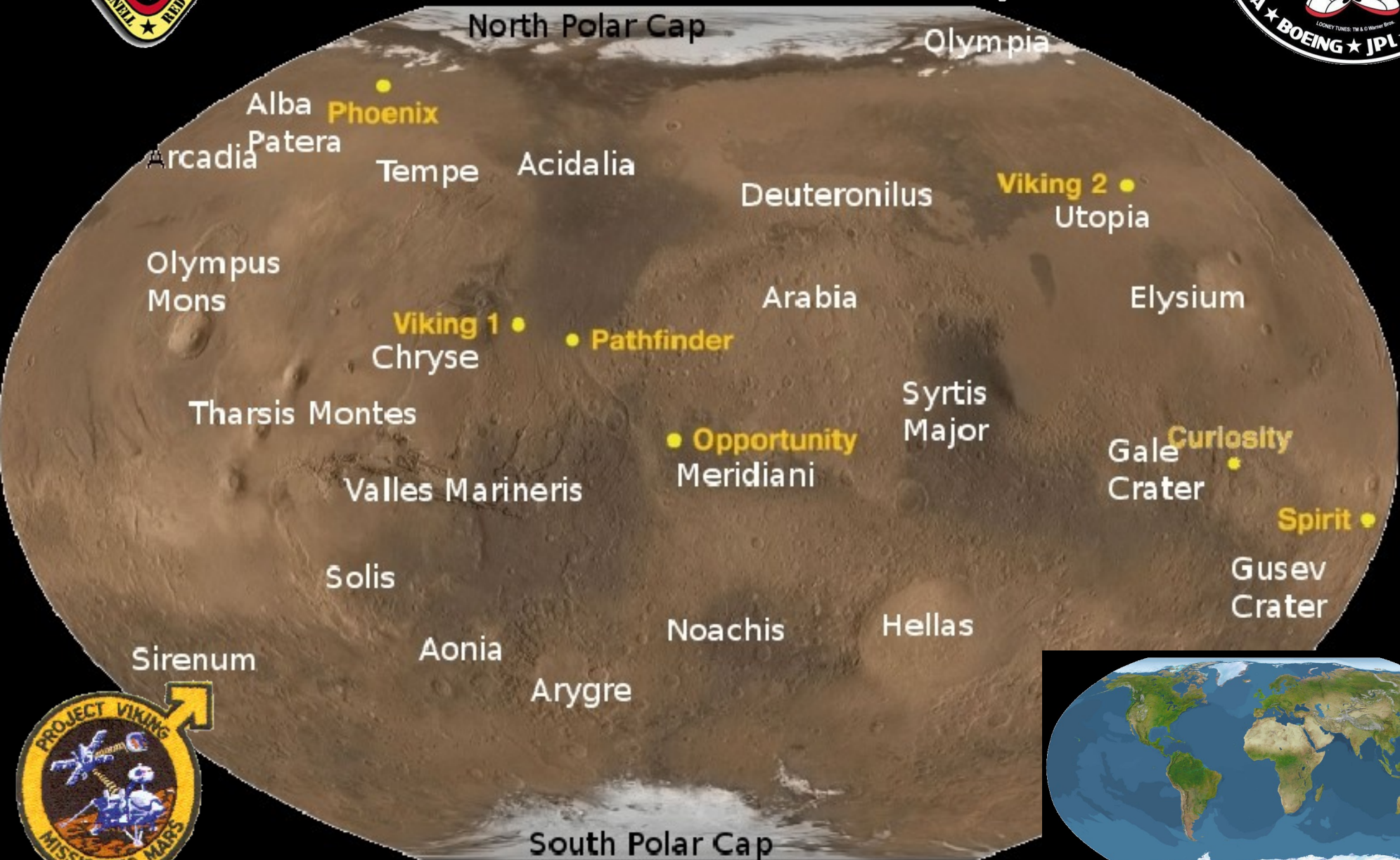


REALLY???

■ Sucesfull ■ Failed



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Martian surface & atmosphere

- * day T below 270 K
- * nights below 190 K
- * Water – no liquid water now, permafrost polar caps, under surface

Imagination

- life?

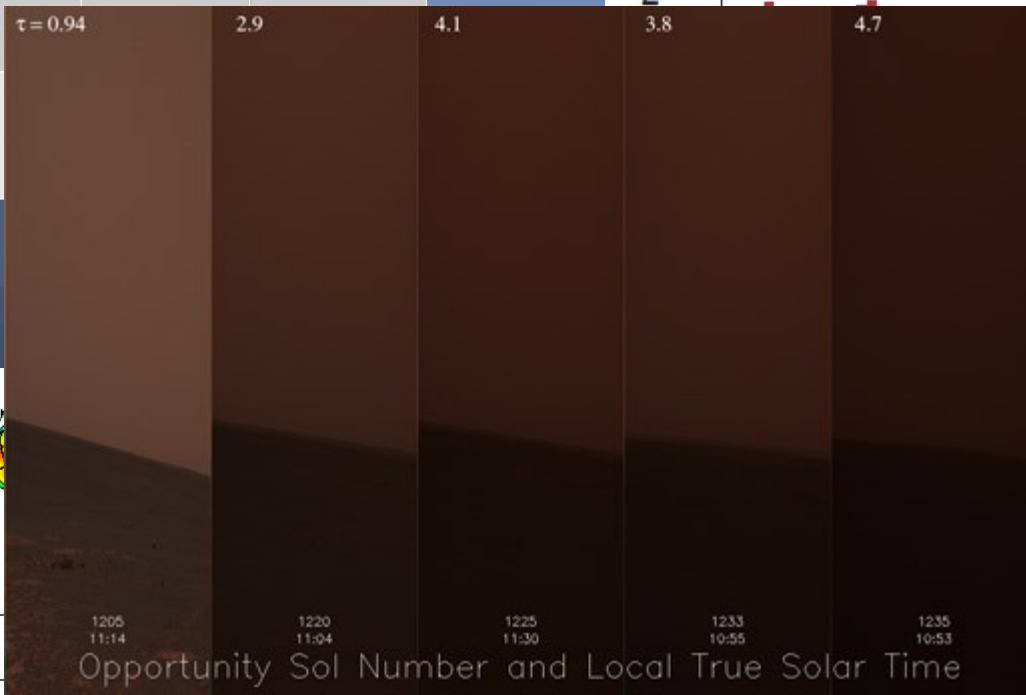
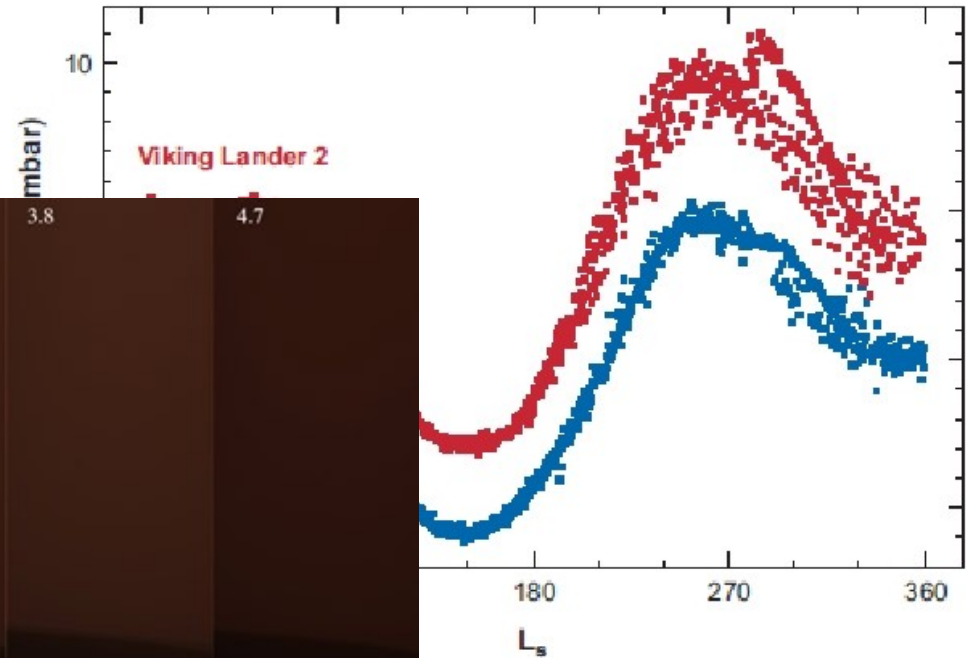
Its origin, the same as on the Earth? exchanging the rock

Atmosphere: thin .. very thin, 7mbar

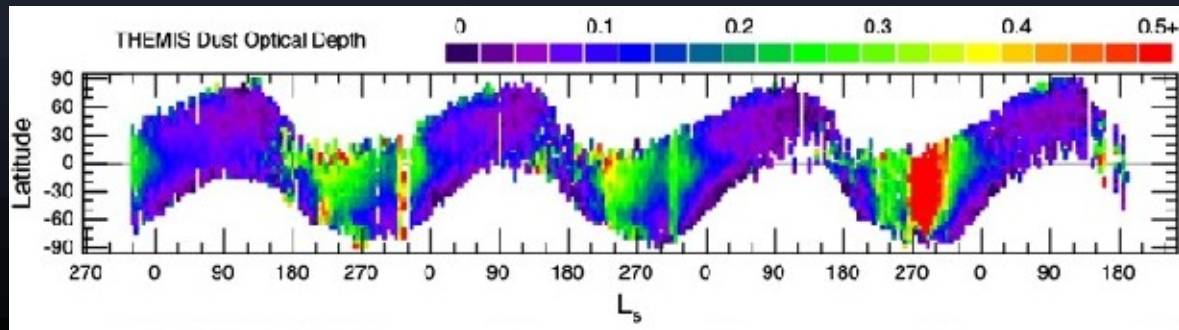
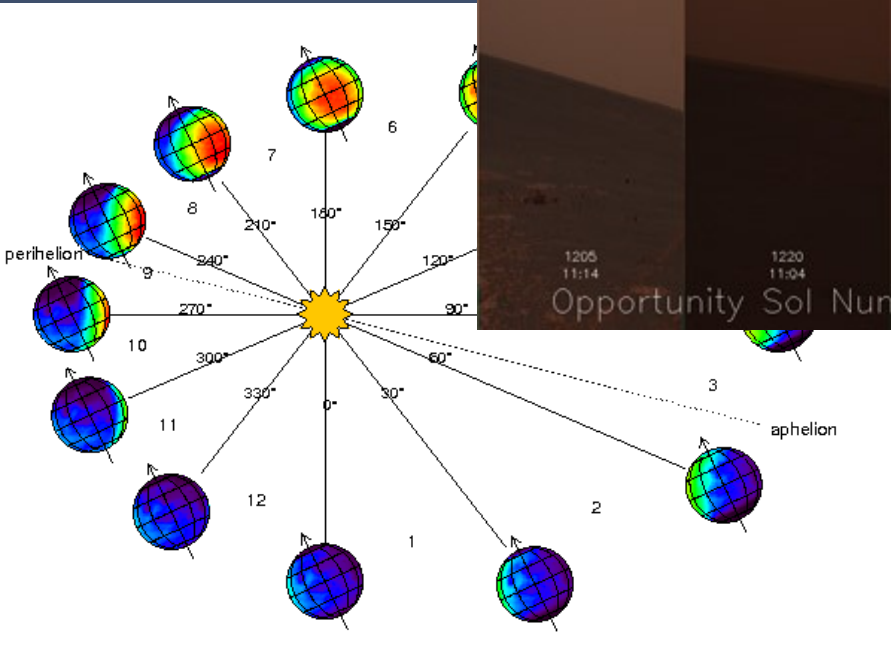
- * no UV protection
- * harsh climate
- * dust, dust activity

Martian atmosphere

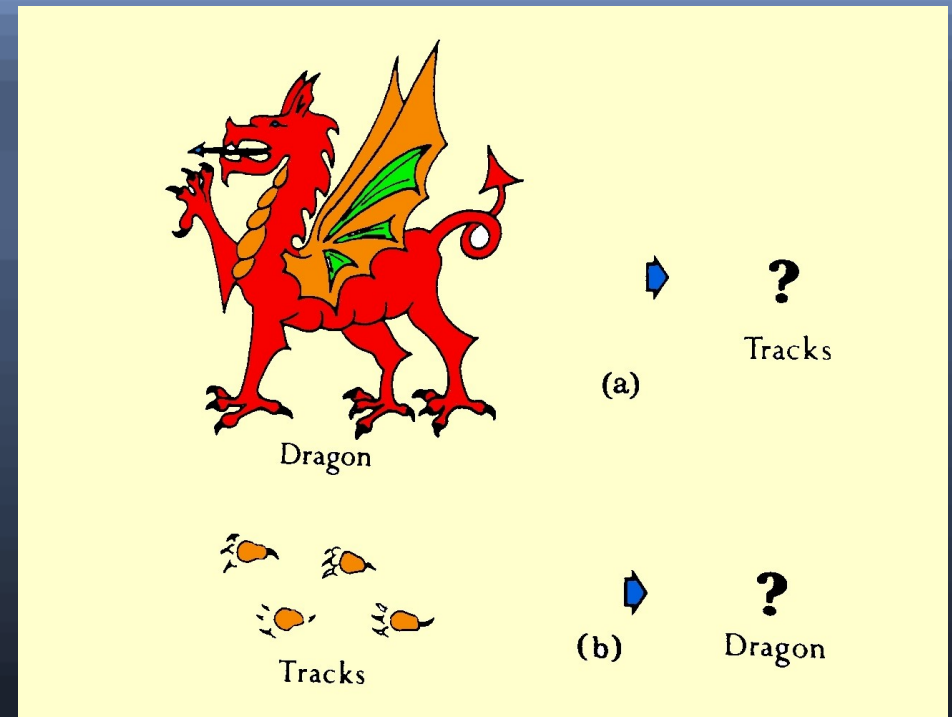
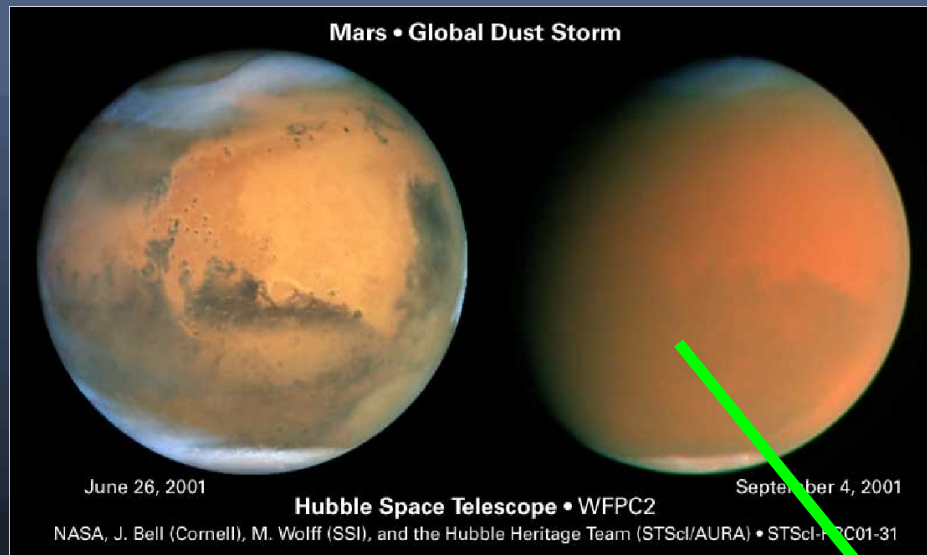
	Venus	Earth	Mars
AU	0.723	$\tau = 0.94$	2.9
Mass [10^{24} kg]	4.87		4.1



GLOBAL STORM!!!!



Dust in Martian atmosphere, what we really know about it?



???

Dust in Martian atmosphere, what we really know about it?

* size (r_{eff} , v_{eff}):

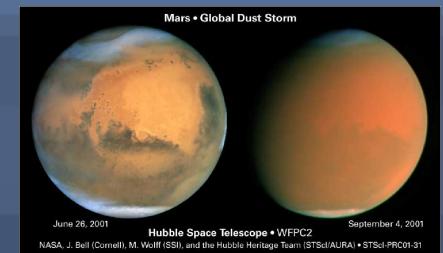
1. conditions
2. retrieval method $r_{\text{eff}} = 0.058 \mu\text{m}$,
 $1.5 \mu\text{m}$,
 $> 9 \mu\text{m}$

* composition:

- MEx atlas
- all rovers coherent chemical results
- Small influence of bedrocks
- Carbonates, e.g. (few percent, Phoenix Lander, Crism on MRO)
- Spectral analogs

* shapes:

- irregular (Phoenix microscope image)
- Modeling the sky brightness changing the shape (size)



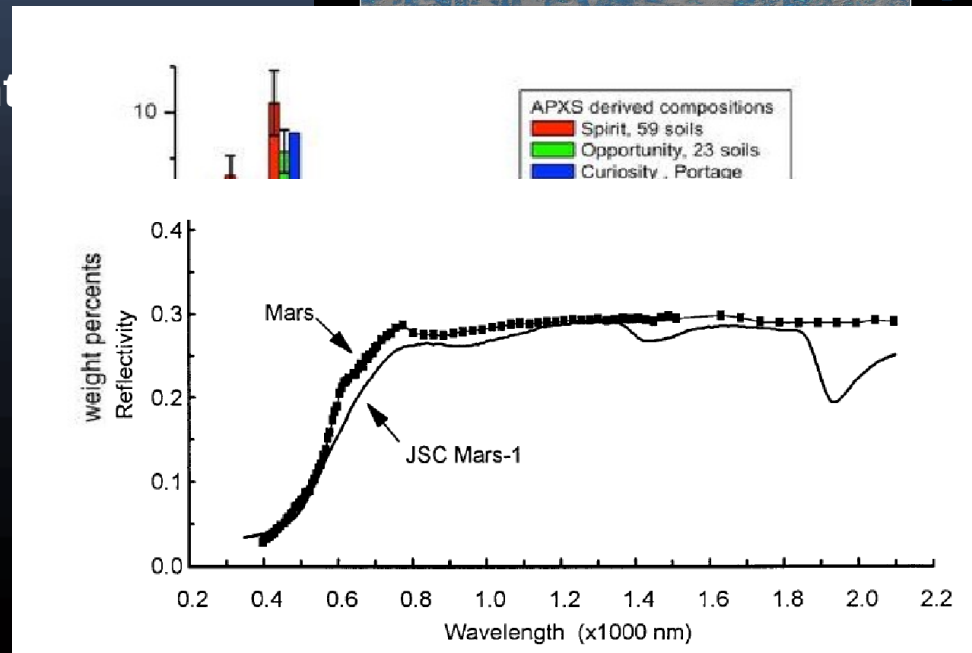
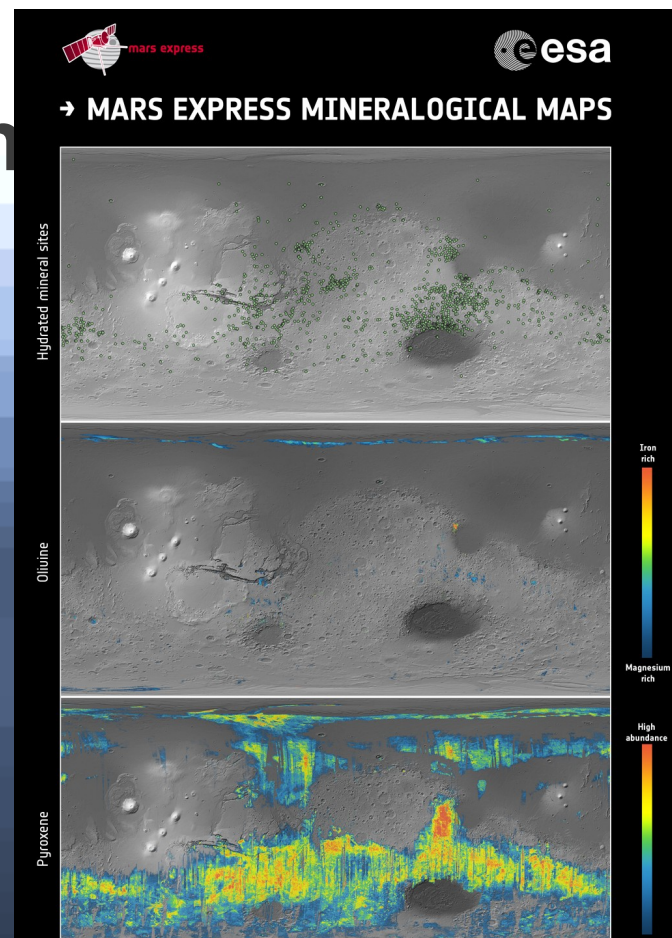
Dust in Martian atmosphere, what do we know about it?

* size (r_{eff} , v_{eff}):

1. conditions
2. retrieval method $r_{\text{eff}} = 0.058 \mu\text{m}$, **$1.5 \mu\text{m}$** , $> 9 \mu\text{m}$

* composition: MEx atlas
 all rovers coherent chemical results
 Small influence of bedrocks
 Carbonates, e.g. (few percent)
 Spectral analogs

* shapes:



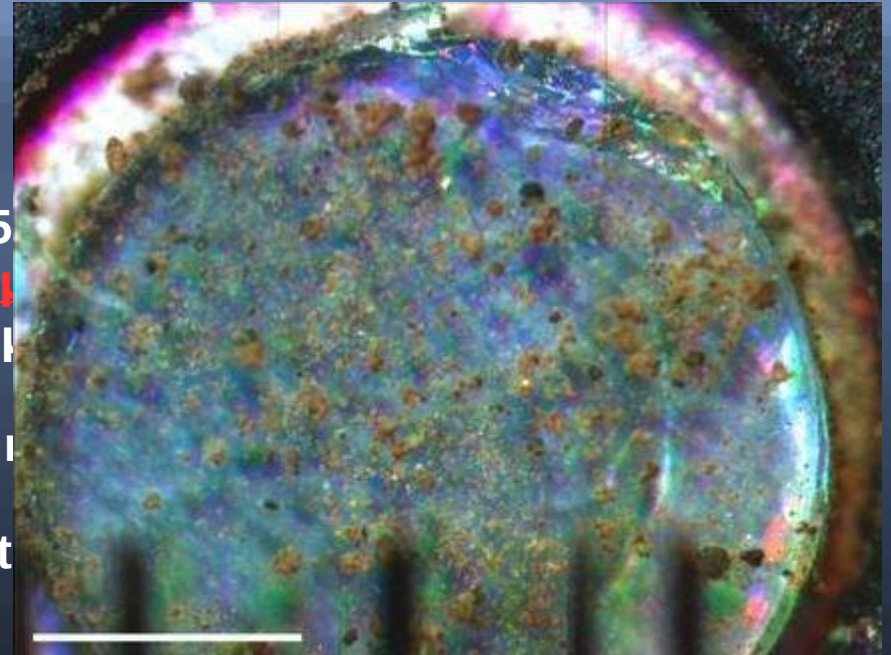
Dust in Martian atmosphere, what we really know about it?

* size (r_{eff} , V_{eff}): 1. conditions
2. retrieval method $r_{\text{eff}} = 0.05$

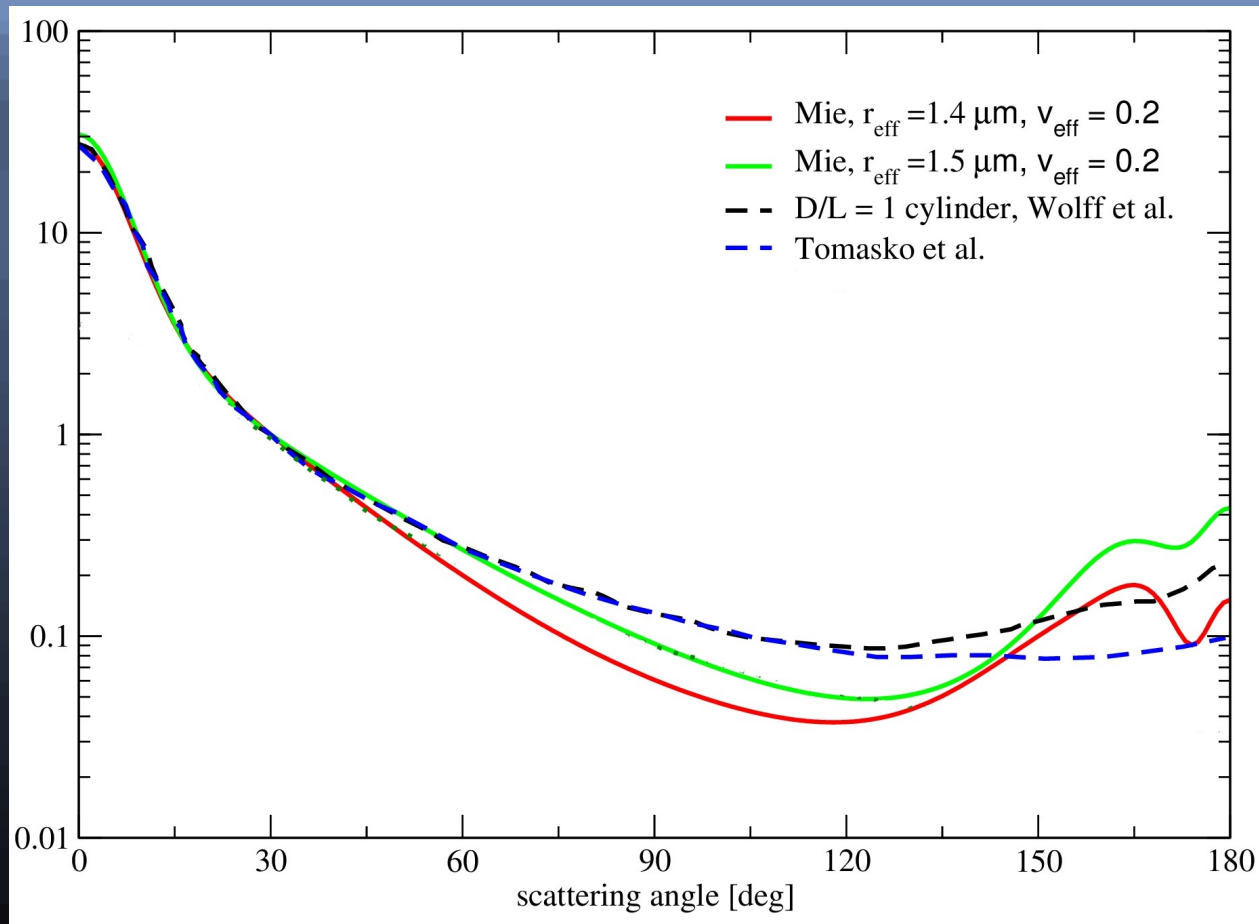
1.5 μm
> 9 μm

* composition: MEx atlas
all rovers coherent chemical
Small influence of bedrocks
Carbonates, e.g. (few percent)
Spectral analogs

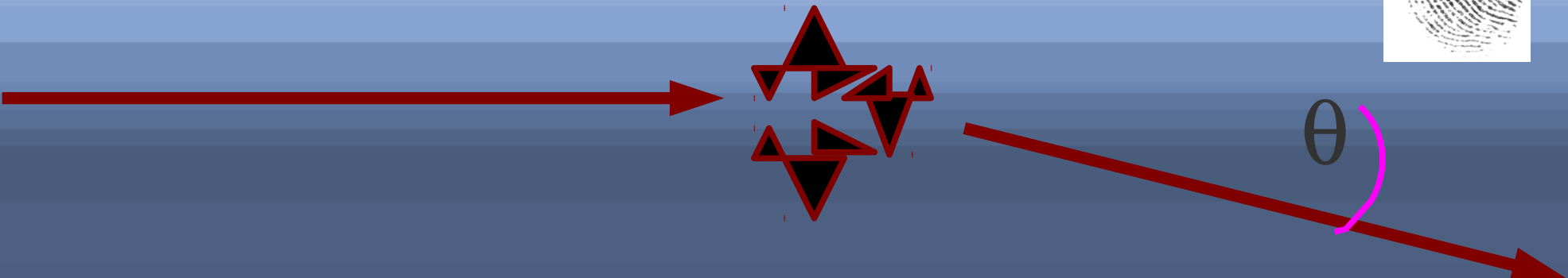
* shapes: irregular (Phoenix microscope image)
Modeling the sky brightness changing the shape (size)



Scattering on Martian dust



Scattering matrix



Intensity

Lin. Polar.

Circ.
Polar.

$$\begin{pmatrix} I_{sc} \\ Q_{sc} \\ U_{sc} \\ V_{sc} \end{pmatrix} \propto \begin{pmatrix} F_{11} & F_{12} & 0 & 0 \\ F_{12} & F_{22} & 0 & 0 \\ 0 & 0 & F_{33} & F_{34} \\ 0 & 0 & -F_{34} & F_{44} \end{pmatrix} \begin{pmatrix} I_{in} \\ Q_{in} \\ U_{in} \\ V_{in} \end{pmatrix}$$

Stokes vector

Scattering Matrix

Stokes vector

Measured scattering matrix (of Martian dust analogs)

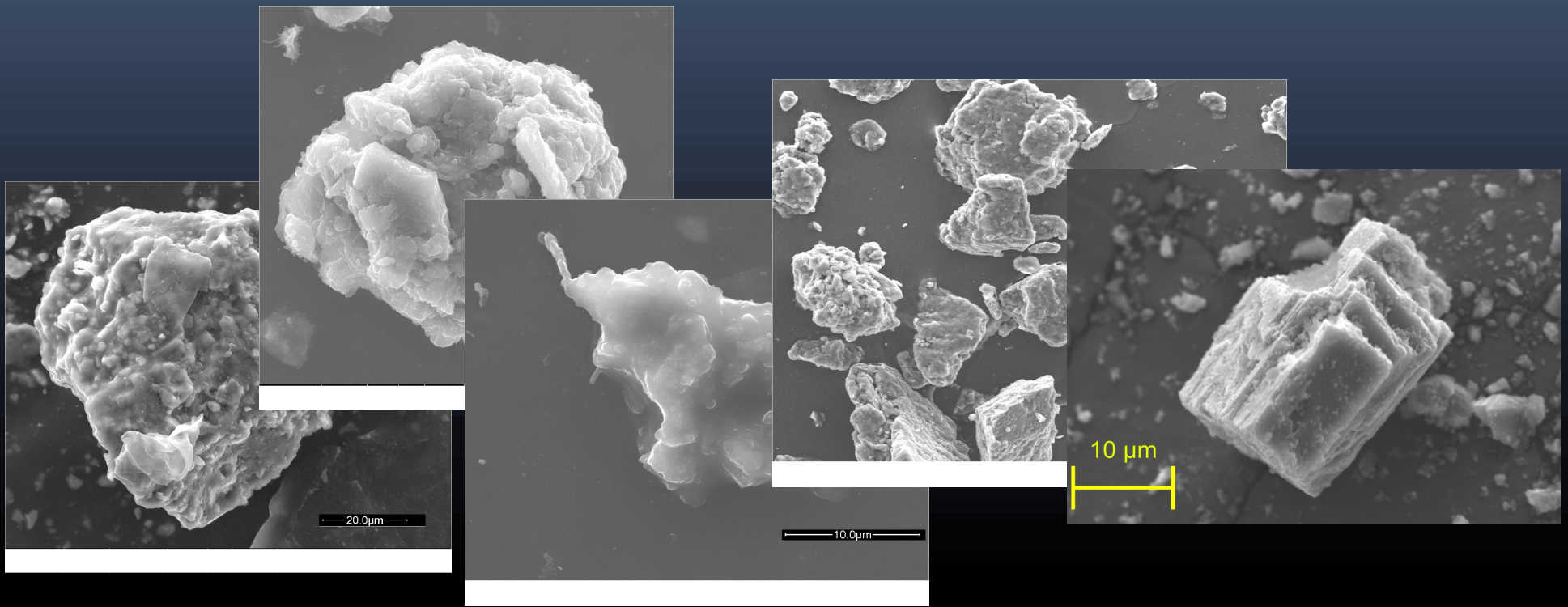
* we have measured 4x4 scattering matrix of:

- Basalt
- JSCs samples (palagonites, one dehydrated)
- Montmorillonite
- Calcite

Past volcanic activity
on Mars

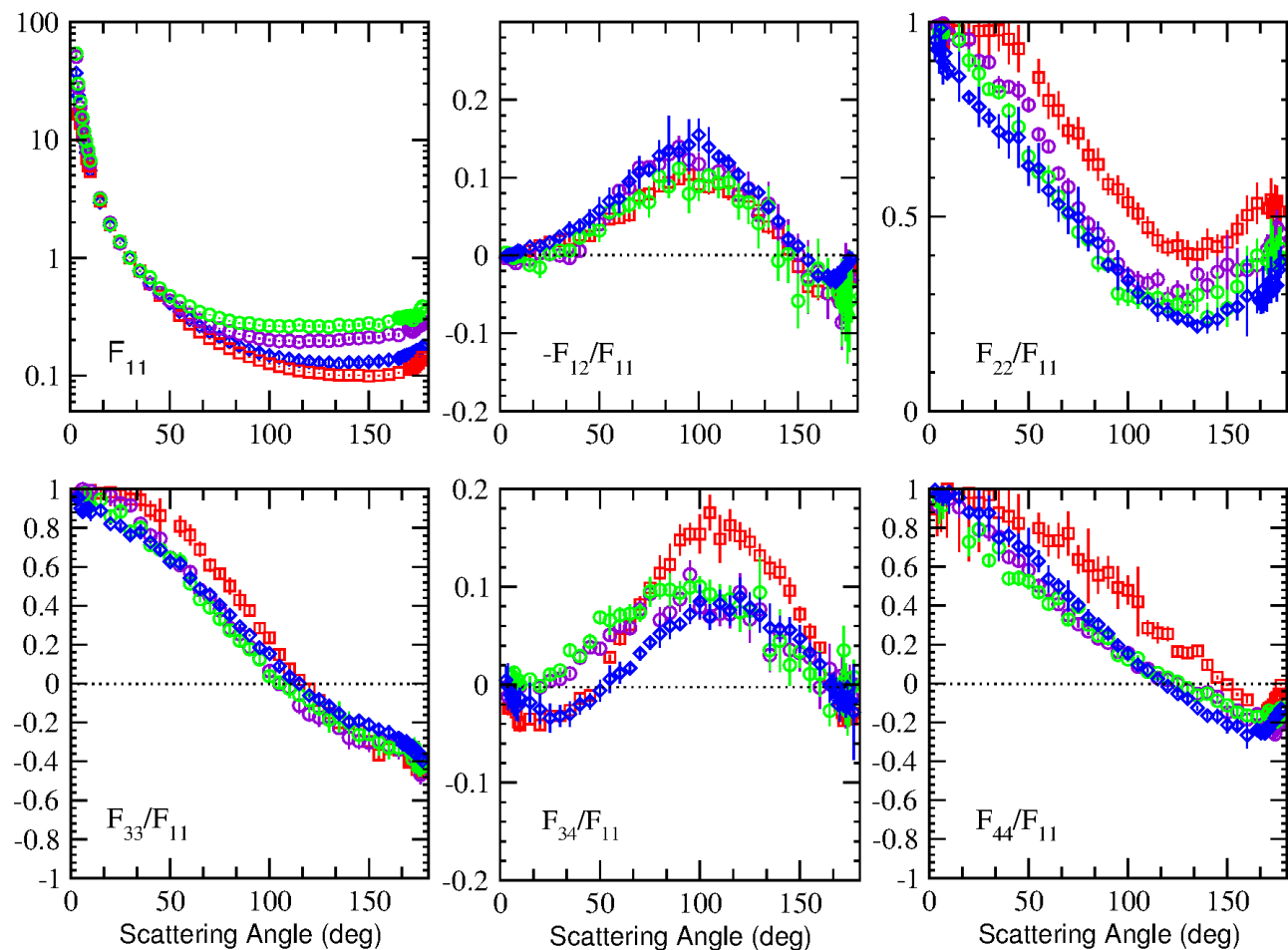
Spectral analogs,
palagonite weathering
product of basalt

Water interest



Measured scattering matrix (of Martian dust analogs). Some results.

- [Calcite](#)
- [Basalt](#)
- [JSC200](#)
- [JSC0](#)



FR:

$R_{\text{eff}}=1.7\mu\text{m}$, $v_{\text{eff}}=7.6$

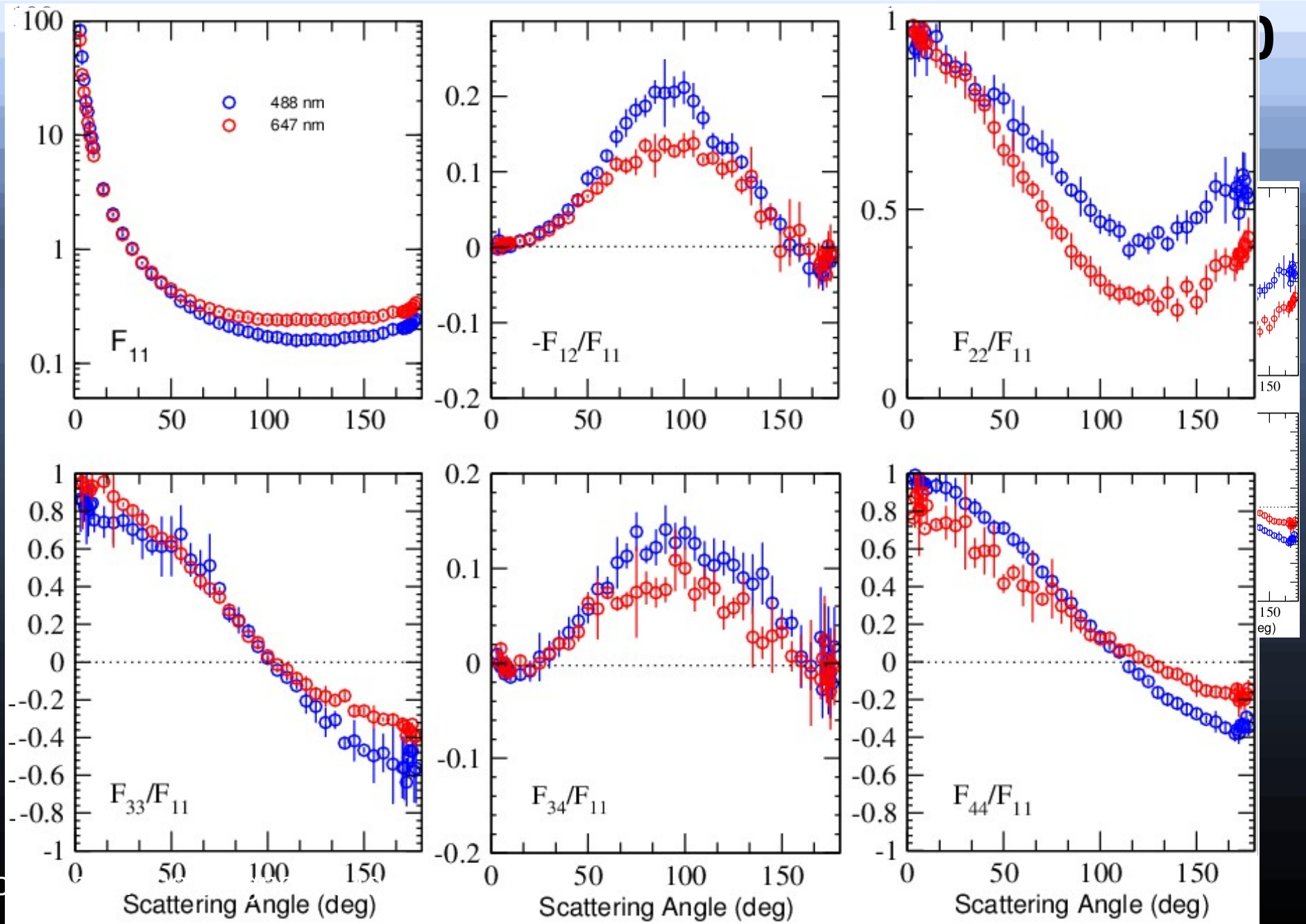
$R_{\text{eff}}=3.0\mu\text{m}$, $v_{\text{eff}}=15.0$

$R_{\text{eff}}=18.2\mu\text{m}$, $v_{\text{eff}}=2.4$

$R_{\text{eff}}=20.3\mu\text{m}$, $v_{\text{eff}}=2.0$

Different RI

Measured scattering matrix



Some conclusions

Scattering matrices database of Martian dust analogs at 2 wavelengths:

- different scattering on regular shaped particles and irregular (better to use our results as radiative transfer models input, not only F11 but entire scattering matrix)
- polarization as diagnostic tool of composition of Martian dust
- much more

